## Homework Problems #1: Fundamentals of Radiation and Fluctuations

- Planck's law of radiation is a triumph of quantum physics and very useful in many fields of science and engineering. (a). Given a 300 K blackbody of unity emissivity, find the frequency in THz (accurate to 100 GHz) and the wavelength in micron where the brightness form of Planck's law is a maximum (can use numerical or analytic techniques). (b) What is this maximum brightness in MKS units ? (c) what is the brightness of this same blackbody at 1.0 THz and how much less is this than the maximum ?
- 2. A useful function for practically any electromagnetic sensor is the *integrated* brightness [i.e., integrated over frequency]. (a) Derive an expression for the integrated brightness IB in terms of the temperature and the Stefan-Boltzman constant,  $\sigma = (2\pi h/15c^2)(\pi k_B T/h)^4$ . (b) Evaluate IB at T = 300 K. (c) Now consider the fractions of IB that lies above and below a certain "cutoff" frequency, IB(>vc) =  $\int_{vc}^{\infty} B(v)dv$  and IB(<vc) =  $\int_{0}^{vc} B(v)dv$  [both are very useful in evaluating the "background" radiation for photon and thermal detectors.]. In the THz region and for terrestrial sources, one can generally apply the Rayleigh-Jeans approximation, hvc/k<sub>B</sub>T << 1. In this case, derive an expression for IB(<vc), and then IB(>vc). (d) Evaluate IB(<vc) for vc = 1 THz and T = 300 K. [suggestion: consult Eisberg&Resnick for an interesting discussion on the genesis of Planck's law and the Stefan-Boltzman constant].
- 3. THz Radiation Filter: (a) Suppose one is using a thermal detector (e.g., Golay cell) that is approximately equally sensitive to IR and THz radiation. Clearly, a low-pass filter is desirable to attenuate (i.e., "stop") the IR but "pass" the THz. For the first approximation, one can assume that the low-pass filter has a power transmission function, T = |S<sub>21</sub>|<sup>2</sup> = τ·θ(ν ν<sub>C</sub>) + θ(ν<sub>C</sub> ν), where θ is the unit step function and τ is the "average" transmission above ν. c. (a) Assuming that the filter is operated at room temperature but does not emit any radiation towards the detector, and that ν<sub>C</sub> satisfies the Rayleigh Jeans limit, write an expression for the value of τ that makes the incident power on the detector from ν >vc equal to that from ν <vc. (b) Evaluate this τ for νc = 1 THz and T = 300, and state the answer as an insertion loss in decibel units. (c) In the stop band region what is a better type of attenuation for the filter can be fabricated as a uniform film that completely intercepts the incoming radiation to the detector ? [clue: there are two reasons for the correct answer, one of which involves Kirchoff's law of radiative transfer].</p>
- 4. Fluctuations of Radiation: the quantum picture. Given the quantum statistics for radiation based on Boltzman and Planck, it is simple to derive the fluctuations from general principles of probability theory.
- (a) Derive the variance or mean-square fluctuations,  $\langle (\Delta n_K)^2 \rangle$  for the radiation modeled as photons having the energy function  $U_K = (n_K + \frac{1}{2})h\omega_K$ , where  $n_K$  is the number of photons in mode K. [clue: start with the Boltzman (exponential) PDF and utilize the general result for random variables,  $\langle (\Delta n_K)^2 \rangle = \langle (n \langle n_K \rangle)^2 \rangle = \langle (n_K)^2 \rangle \langle n_K \rangle^2$ ].

(b) Use the above expression to calculate the root mean square fluctuations of incident power in a single spatial mode from the sun within a spectral bandwidth of 1 GHz at the following two center wavelengths: (1)  $\lambda = 0.5$  mm (THz region), and (2)  $\lambda = 0.5$  micron (visible region). Express your answer in MKSA and in dBm (decibels relative to 1 mW). [clues: make the narrow passband approximation in both cases; approximate the sun as having a brightness temperature of 5800 K].