

- Notes Set 11: Noise of miscellaneous devices

- Schottky and PN junction diodes. transmission lines. antennas
- antenna noise temperature. Temperature of the sky. Friss formula in temperature form.

ECE Notes set II

We have now developed noise models of:

\* bipolar transistors

\* Field-effect transistors.

\* Resistors

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We must also consider noise models of

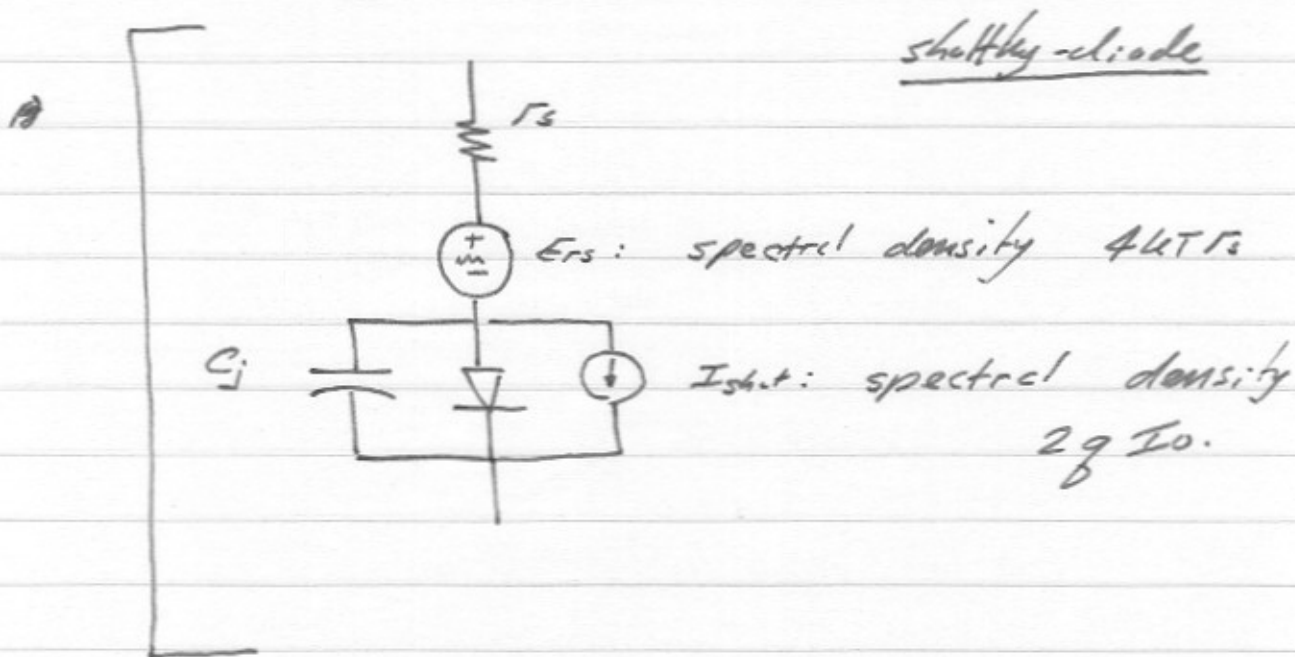
\* diodes

\* Antennas

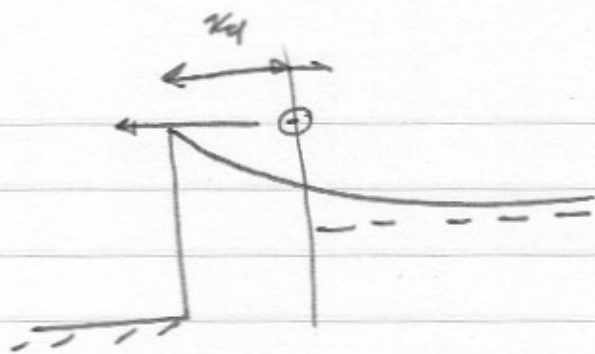
\* Transmission Media:

Diodes:

We have largely covered these, but a few corrections remain: The noise current has a spectral density  $2q(I_{dc} + 2I_0)$ , but



real diodes have series resistance & junction capacitance which must be added to the model.

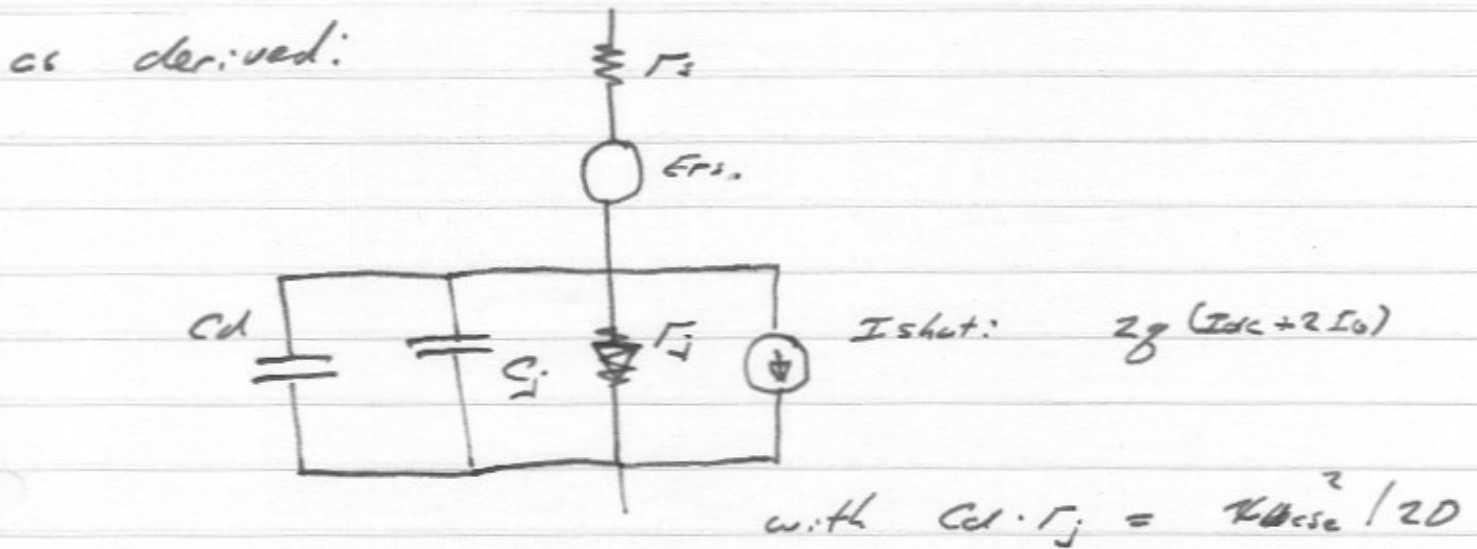


Schottky diodes have a depletion region of width  $x_d$ , giving a transit time in forward bias of  $T_T \approx K x_d / v_{thermal}$ \*, which will change the shot noise spectral density to

$$2q I_0 \cdot \left[ \frac{\sin \omega T_c}{\omega T_c} \right]^2 \approx \frac{2q I_0}{1 + \omega^2 T_c^2}$$

\* this can become a long discussion! Avoid it.

For P.N. junction diodes in the short-base limit, and for thin depletion regions, the diode model includes a diffusion capacitance as derived:



\* For long-base diodes  $r_j$  &  $C_d$  become frequency-dependent, as does  $I_{shot}$ . This was given earlier.

\* For long depletion regions (P-I-N) diodes, the device analysis becomes lengthy whether for noise or just small-signal. See Sze.

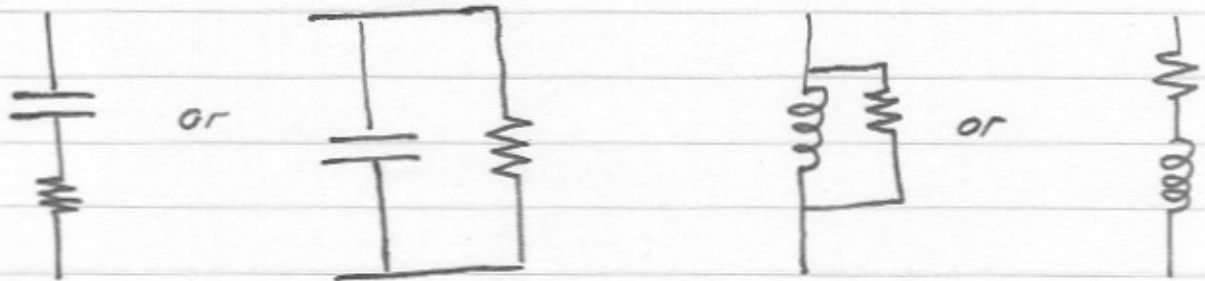


What about Inductors, Capacitors, and transmission lines?

Ideal  $\infty\Omega$ ,  $\frac{1}{T}$ , and transmission lines have ~~no~~ no means of dissipating power. Their terminal Impedance is pure imaginary, so the available noise power is zero.

Real Inductors or Capacitors have loss.

Regardless of the physical origin of this loss, it can be modelled by:



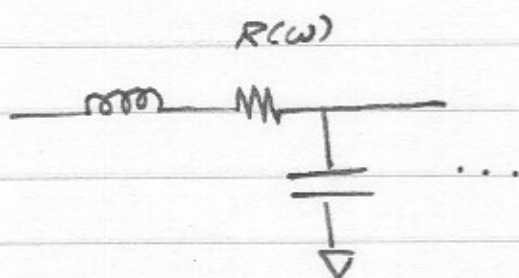
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... where the series or shunt resistance may be frequency-dependent. In all cases that resistance has thermal noise with available power  $kT$ , where  $T$  is the local temperature.

## Transmission-Lines (wires)

require a bit more attention. 3 causes of loss

### 1) Conductor resistance



$R(\omega)$  is approximately of the form  $R_{dc} + K\sqrt{f}$

where the latter term is due to the skin effect.

$R(\omega)$  can be simply treated as any other resistor, & has available noise power  $KT$ .



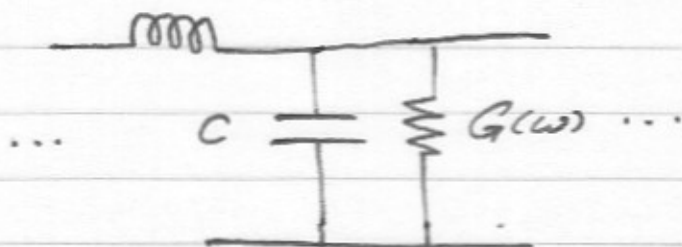
2) Dielectric losses:

$$\epsilon = \epsilon_r \epsilon_0 + j \epsilon_{ir} \epsilon_0$$

$\epsilon_{ir} / \epsilon_0$  is called the loss-tangent.

... This is power lost in polarizing & de-polarizing the dielectric..

usually modelled as such:



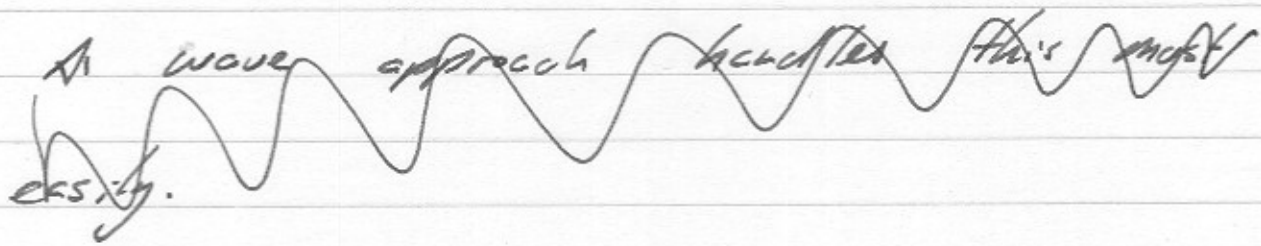
where  $G / j\omega C = \epsilon_{ir} / j\epsilon_r$

and  $G$  has available power  $kT$ .

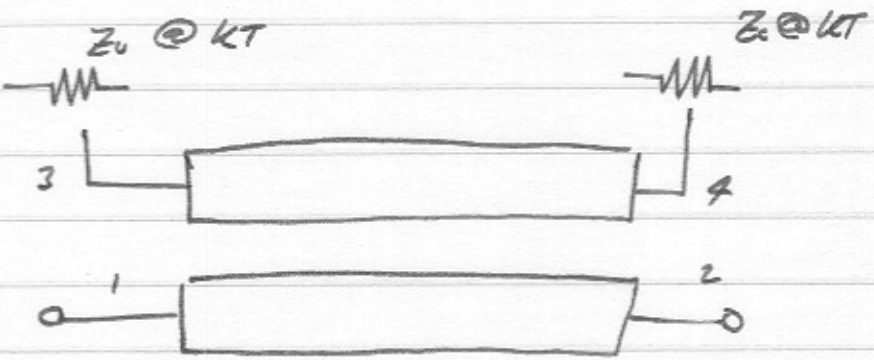
Fortunately  $G(\omega)$  is usually negligible compared to skin loss.

3) open transmission lines have radiation losses if completely open, or "mode coupling losses" if the open line is contained within a larger closed package. This is radiation, and the line acts as a transmitting antenna. By reciprocity the line also acts as a receiving antenna, and generates thermal noise power from the thermal radiation of the local environment.

An wave approach handles this effect  
density.



It is tempting to model the transmission line for noise thus:

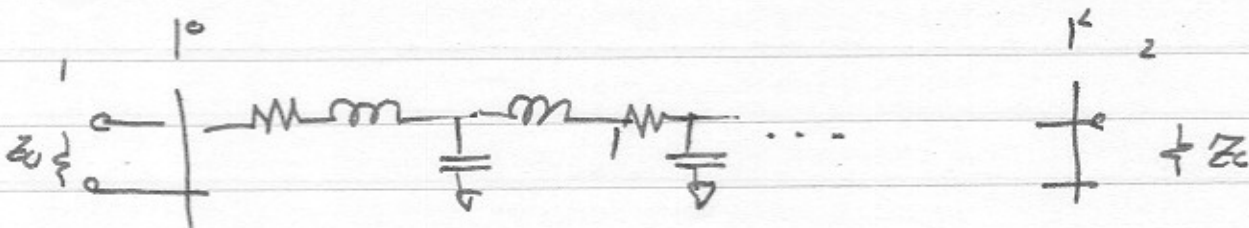


lossless directional coupler

such that all ports have impedance  $Z_0$  & the directional coupler is prebed so that the port 1-2 coupling is equal to the line loss.

But the sources at ports 3 & 4 are not generally uncorrelated

This can be seen for skin loss:



If we call  $E_{in}$  and  $E_{out}$  the voltage wave

amplitudes coupling to ports 1 and 2

$$E_{in} := \int_0^L E_n(z) e^{-j\beta z} e^{-\alpha z} dz$$

$$E_{out} := \int_0^L -E_n(z) e^{-j\beta(L-z)} e^{-\alpha(L-z)} dz$$

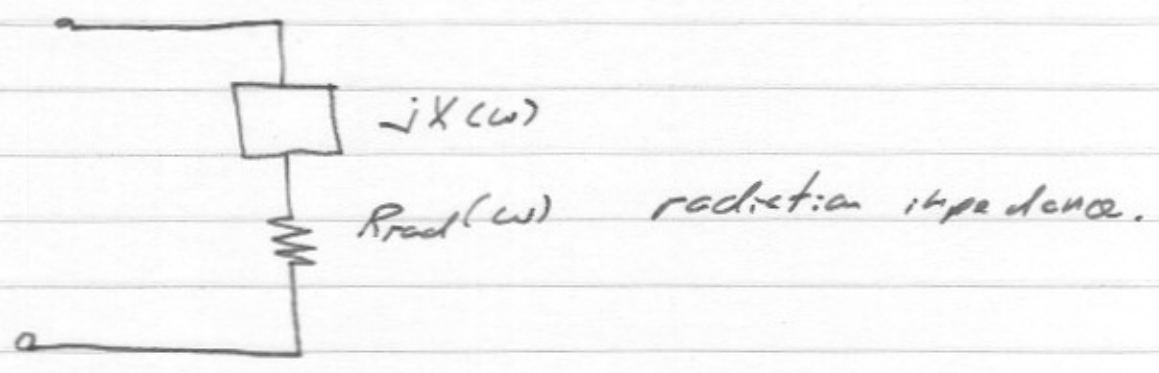
where  $E_n(z)$  has spectral density  $4kT R(z)$ .

...  $E_{in}$  and  $E_{out}$  are clearly correlated!

# Antenna Noise

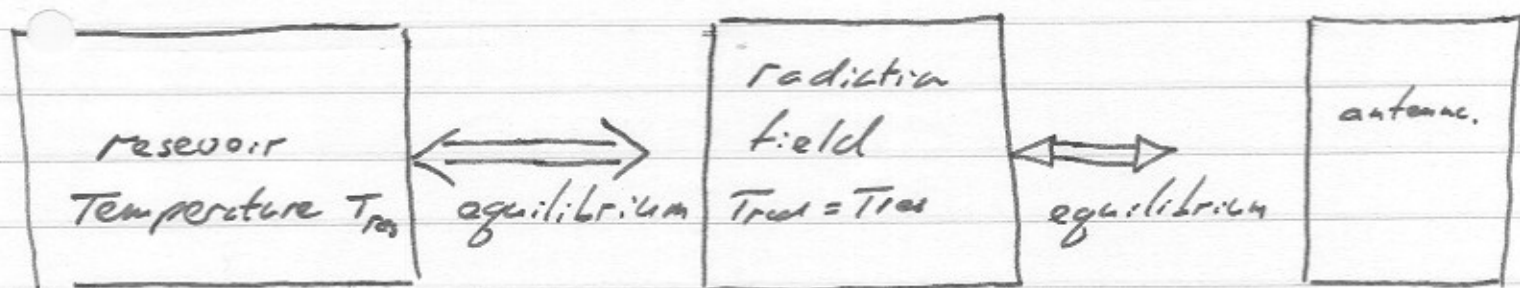
The terminal impedance of an antenna

is modeled as so:



If we couple the antenna to a radiation field, we can compute its available noise power:





So, the antenna has available noise power

$$\approx \underline{\underline{KT_{rad}}}. \quad \Rightarrow \quad \left[ \frac{\partial \langle P_{cu} \rangle}{\partial f} = \frac{1}{2} hf + \frac{hf}{\exp(hf/kT) - 1} \right]$$

Note that  $T_{rad}$  is the temperature of the radiation field to which the antenna is coupled.

This depends very much on where the antenna is aimed.

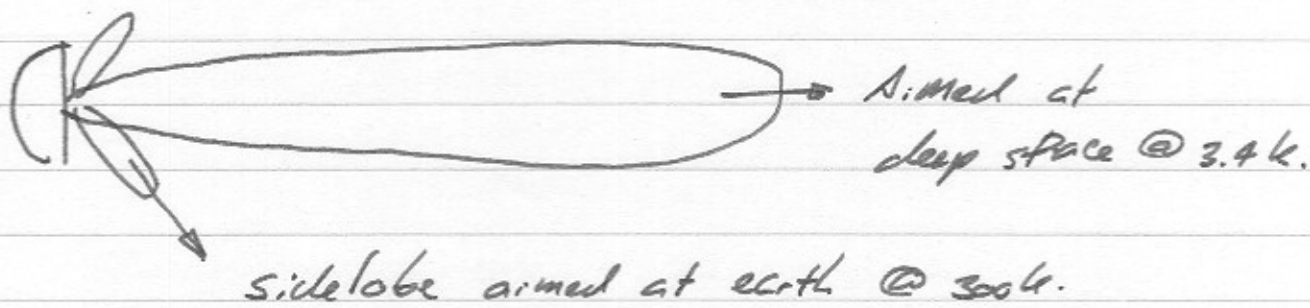
Aimed at deep space:

Cosmic background radiation: 3.4 K.

Aimed near planet (earth's) surface: 300 K:

Antennas for radio astronomy must be designed

for very low sidelobes:



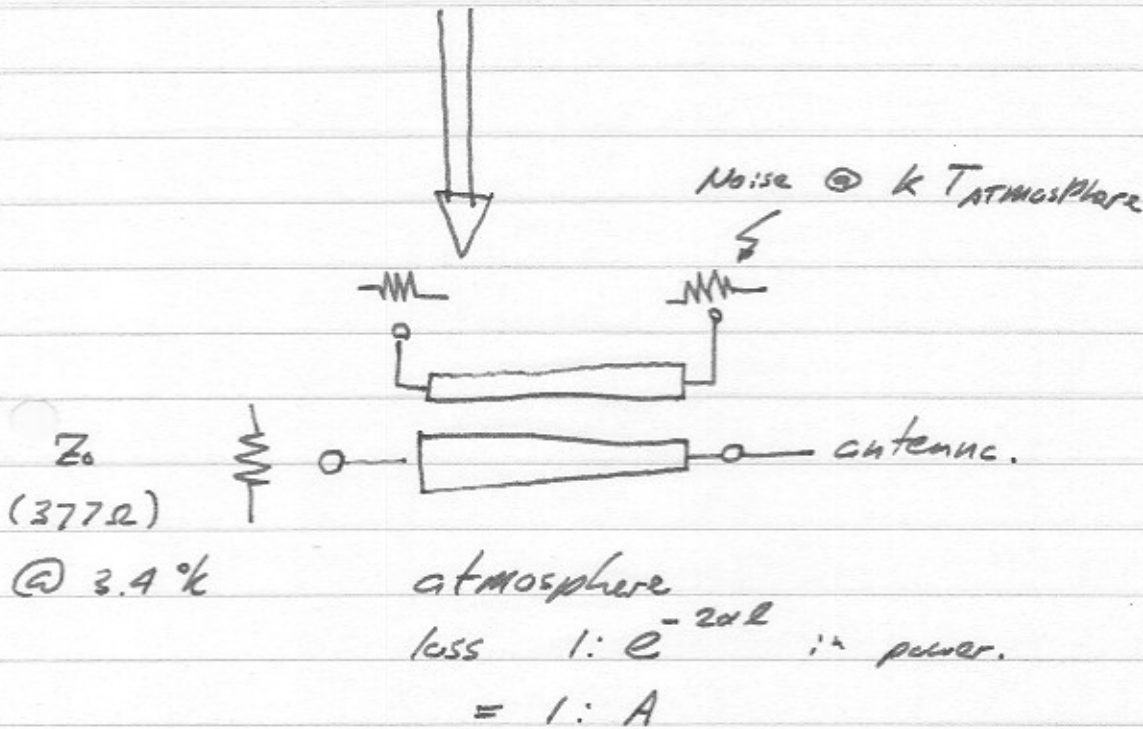
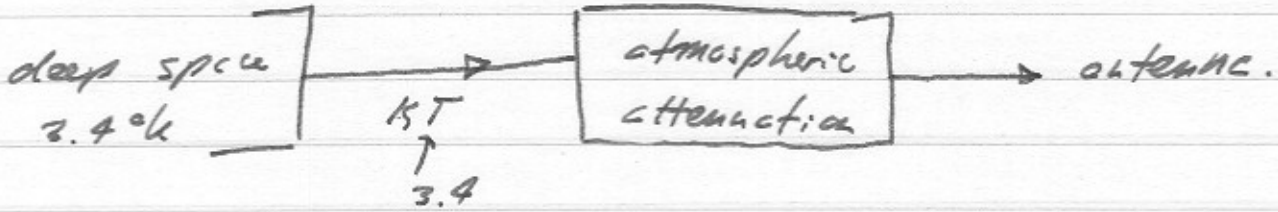
[A 1% sidelobe aimed at earth generates as much noise power as the target!]

Antennas for Astronomy are made as big as much

for sidelobe noise suppression as they are for gain

or directivity

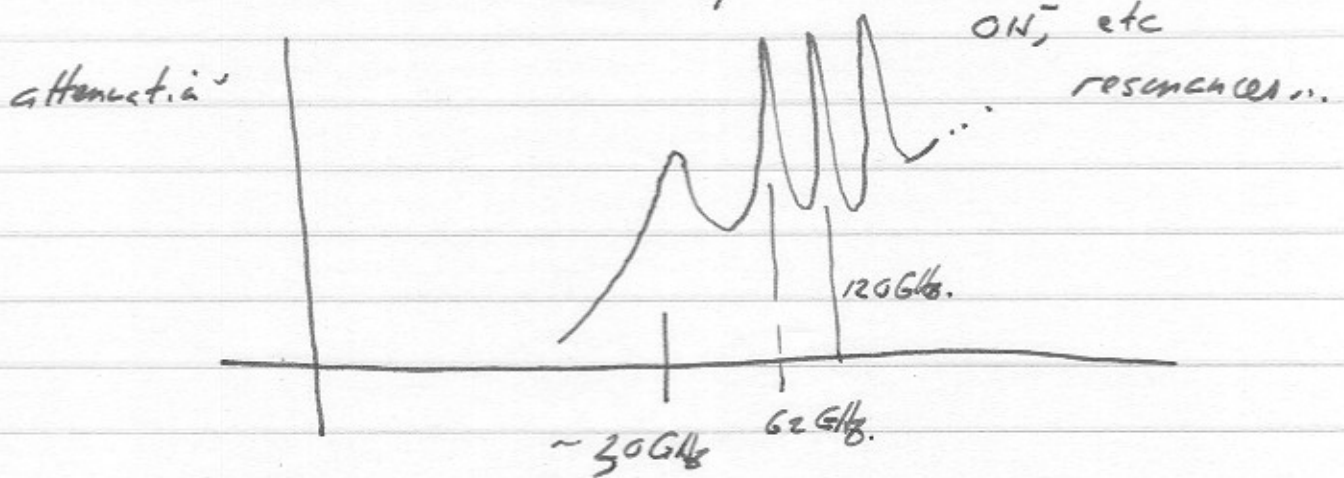
Atmospheric Effects can be modelled as an attenuator:



the antenna will see an available noise power

$$T_{red} \cdot (A) + T_{ATMOSPHERE} \cdot (1-A)$$

The atmosphere is at  $250-300^{\circ}\text{K}$ , and has an attenuation curve something like this



So an antenna pointed at the Zenith would produce strong noise peaks at 30, 62, & 120 GHz.



Real antennas also have conductor resistance, from which the model becomes:



As deep space / satellite antennas see  $T_{rad} \ll 300\text{K}$ , it is imperative to obtain  $R_{loss} \ll R_{rad}$ .

Note that terrestrial communications systems will see  $T_{rad} \sim 300\text{K}$ , whilst satellite receivers ~~see~~ may see  $T_{rad} \sim 10 - 50\text{K}$ . Requirements will differ greatly.



Note that noise figure:

$$F \triangleq \frac{S/N|_{IN}}{S/N|_{out}} = \frac{kT + N_{amp}}{kT}$$

... where  $N_{amp}$  is the amplifier input-referred noise power contribution.  $F$  is normally specified for a 300 K background for  $kT$ , yet an amplifier may be used with an antenna having  $T_{rad} \ll 300$  K.

In this case,  $F=1.1$  vs  $F=1.01$  can represent vastly different % noise contributions from the amplifier. Better to either use  $F$  defined at the antenna noise radiation noise temperature or use the equivalent noise temperature:

$$\boxed{T_{eq} = N_{amp} / k.}$$

From which:

$$T_{eq} = [F(T_0) - 1] \cdot T_0$$

where  $F(T_0)$  is the noise figure at reference temperature  $T_0$

or

$$F(f) = 1 + \frac{T_{eq}}{T_0}$$

Note that noise temperature is not physical  
 a 10dB 50Ω attenuator at 300K room temperature has a 3000K noise temperature.  
 There is nothing that hot in the thermodynamic sense.

We can write the Friis Noise formula in terms of noise temperatures:

$$F_T = F_1 + \frac{F_2 - 1}{G_{cu1}} + \frac{F_3 - 1}{G_{cu1} G_{cu2}} + \dots$$

$$T_0 + T_{eqT} = T_0 + T_{eq1} + \frac{T_{eq2}}{G_{cu1}} + \frac{T_{eq3}}{G_{cu1} G_{cu2}} + \dots$$

$$T_{eqT} = T_{eq1} + \frac{T_{eq2}}{G_{cu1}} + \frac{T_{eq3}}{G_{cu1} G_{cu2}} + \dots$$

The  $T_{eq}$ 's are calculated for each device assuming the output impedance of the prior stage.

This should remind you how little  $T_{eq}$  has to do with a physical temperature  $\left[ \frac{1}{T} \stackrel{\Delta}{=} \frac{\partial S}{\partial E} \right]$